(3) Today Next Class (4)

- 2.2 The Schrödinger Equation
- 2.2.1: The Particle in a Box
- 2.2.2 Quantum Numbers and Atomic Wave Functions
- 2.2.3 The Aufbau Principle and Shielding

### (5) Second Class from Today

- 2.2.2 Quantum Numbers and Atomic Wave Functions
- 2.2.3 The Aufbau Principle and Shielding
- 2.3 Periodic Properties

2.2.2 Quantum Numbers and Atomic Wave Functions

2.2.3 The Aufbau Principle and Shielding

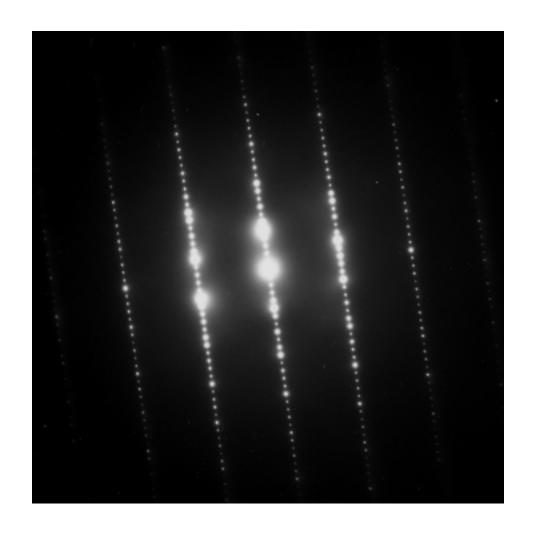
2.3 Periodic Properties

## **Third Class from Today** (6)

3.1 Lewis Structures

3.2 VSEPR

## Wave-Particle Duality



de Broglie  $\lambda = h/mv$ 

Heisenberg

$$\Delta x \, \Delta p_x \ge h/4\pi$$

Wave Mechanics and The Schrödinger Equation			J 41	Section 2.2
Hamiltonian				•
—→ Н <b>Ч</b>	= EΨ (linear	- combina	tions o	I the x
	ofal	1 the e-1	5)	

Squaring the wave function gives use the probability of finding a the electron at a given location in space

The wave function must be an eigenfunction

wave functions are cho-	ren bare on mathematical		
Necessary qualities of a wave function			
1. The wave function must be single valued.	Cannot have two probabilities for finding		
	the electron at a given point		
2. The wave function and its first derivatives must	The probability must be defined at all		
be continuous.	points in space and cannot change abruptly		
3. The wave function must approach 0 as r	The probability must get smaller at large distances of the atom. The atom must be		
approaches infinity	finite.		
4. Integrating $\Psi_A\Psi_A^*$ over all space must equal 1	The electron must be somewhere in space.		
	Process is called normalizing the wave		
	function		
5. Integrating $\Psi_A\Psi_B^*$ over all space must equal 0	The orbitals must be orthogonal (mutually exclusive)		

$$H = \frac{-h^2}{8\pi^2 m} \left( \frac{\delta^2}{\delta x^2} + \frac{\delta^2}{\delta y^2} + \frac{\delta^2}{\delta z^2} \right) - \frac{Ze^2}{4\pi\epsilon_0 \sqrt{x^2 + y^2 + z^2}}$$

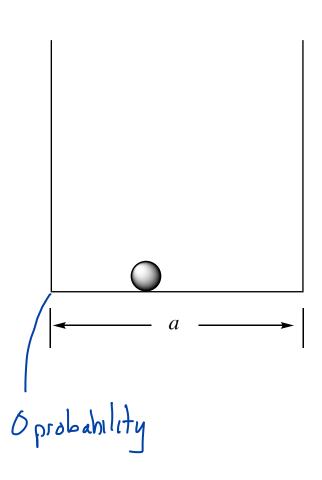
Since,

$$E = \frac{1}{2} m v^{2} + \frac{k^{2}q_{1}q_{2}}{r}$$

$$r = \sqrt{x^{2} + y^{2} + z^{2}}$$

$$KE$$

$$H = \frac{-h^{2}}{8\pi^{2}m} \left(\frac{\delta^{2}}{\delta x^{2}} + \frac{\delta^{2}}{\delta y^{2}} + \frac{\delta^{2}}{\delta z^{2}}\right) - \frac{Ze^{2}}{4\pi\epsilon_{0} r}$$



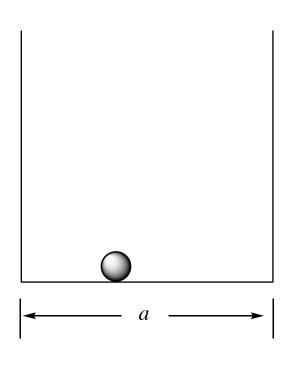
$$\Psi = A \sin rx + B \cos sx$$

$$problem ... (f x = 0)$$

$$\cos s(o) = 1$$

$$B \text{ must be } 0$$

$$\Psi = A \sin rx$$

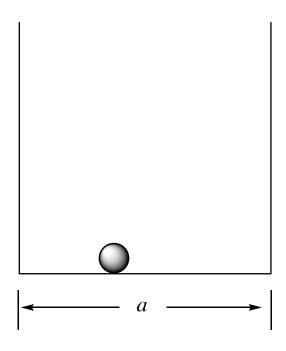


$$\Psi = A \sin rx$$

$$H = \frac{-h^2}{8\pi^2 m} \left( \frac{\delta^2}{\delta x^2} + \frac{\delta^2}{\delta y^2} + \frac{\delta^2}{\delta z^2} \right) - \frac{Ze^2}{4\pi\epsilon_0 r}$$

$$H \quad \forall \qquad E \quad \psi$$

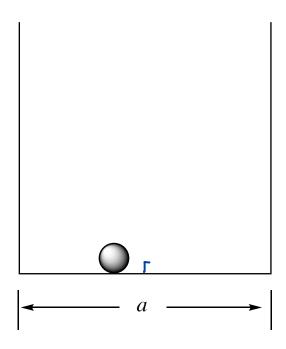
$$\frac{-h^2}{8\pi^2 m} \left( \frac{\delta^2}{\delta x^2} \left( A \sin rx \right) \right) = E(A \sin rx)$$



$$\frac{-h^2}{8\pi^2 m} \left( \frac{\delta^2}{\delta x^2} \left( A \sin rx \right) \right) = E \left( A \sin rx \right)$$

$$\frac{-h^2}{8\pi^2 m} (Ar) \left( \frac{\delta}{\delta x} \left( \cos rx \right) \right) = E \left( A \sin rx \right)$$

$$\frac{-h^2}{8\pi^2 m} \text{ (-Ar}^2\text{) (sin rx)} = \text{E A sin rx}$$

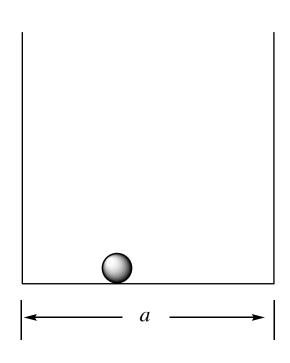


$$\frac{-h^2}{8\pi^2 m} (-Ar^2) (\sin rx) = E A \sin rx$$

$$\frac{-h^2}{8\pi^2 m}$$
 (-r<sup>2</sup>) = E

$$r^2 = E \frac{8\pi^2 m}{h^2}$$

$$r = \frac{2\pi}{h} \sqrt{2mE}$$



$$r = \frac{2\pi}{h} \sqrt{2mE}$$

But remember

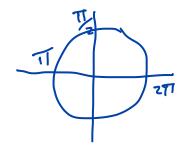
 $\Psi = A \sin rx$ 

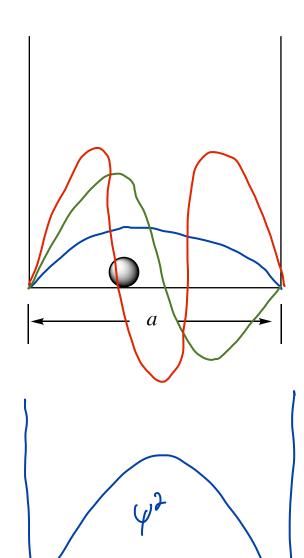
r must also equal

So,

 $n\frac{\pi}{a}$  counting # T = a = T

$$\mathsf{E} = \frac{n^2 h^2}{(8a^2\mathsf{m})^2}$$





$$\Psi = A \sin rx$$

$$r = n \frac{\pi}{a}$$

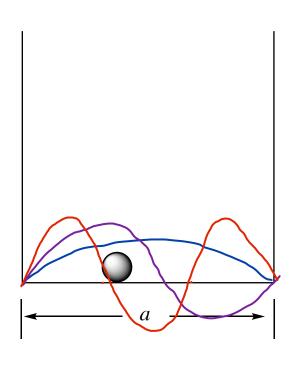
$$\Psi = A \sin(n \frac{\pi}{a} x)$$

normalize the wave function

As you = 1 in this case," all space"

1s from x=0 to x=a

lowest energy state are are most likely to find the ball in the middle of the box



$$\Psi = A \sin(n \frac{\pi}{a} x)$$

$$(\Psi\Psi^*) = 1$$

$$\Psi = (2/a)^{1/2} \sin (n\pi/a) x$$

excited states ground stade for particle
In its lowest energy stade
are are most likely to find
the ball in the middle of
the box Results - Section 2.2

# Equations

https://www.westfield.ma.edu/cmasi/advinorg/angular distribution functions/ text and graphics containe.htm

### **Pictures**

https://www.westfield.ma.edu/cmasi/advinorg/quant\_orbital\_surfaces/orbital\_surfaces.htm

Models

s and p

https://www.westfield.ma.edu/cmasi/organic/mo-plain/aos.html

d orbitals

https://www.westfield.ma.edu/cmasi/advinorg/dorbs/dorbsp.html

One quantum number wasn't enough to model the electrons in an atom

n is the principal quantum number

allowed values are the whole numbers starting a 1

*I* is the Angular momentum quantum number

N-1

allowed values are the from to 0 in whole number increments

m<sub>I</sub> is the magnetic quantum number

allowed values are the from -1 to I in whole number increments

ms is the spin quantum number

allowed values are  $-\frac{1}{2}$  or  $+\frac{1}{2}$ 

Orbitals (n, l, and m<sub>l</sub>) Section 2.2.2

## The angular distribution function:

The square of the angular distribution function describes the probability of finding the electron at angles  $\theta$  and  $\phi$ . In other words, picture yourself standing on the nucleus, and the electrons look like fog around you. The square of the angular distribution function describes how thick fog is when you look in different directions.

The Angular distribution function describes the basic shape of the orbital, or the number of nodal planes in an orbital.

The angular distribution functions depend only on the quantum numbers l and  $\mathbf{m}_l$ . That is, the angular distribution functions of all electrons with the same l and  $\mathbf{m}_l$  values are the same regardless of  $\mathbf{n}$  or  $\mathbf{m}_s$ . Another way of saying this is that all s orbitals have the same basic shape. For example The 2s orbital (n = 2, l = 0,  $m_l = 0$ ), the 3s (n = 3, l = 0,  $m_l = 0$ ) and the 4s (n = 1, l = 0,  $m_l = 0$ ) have the same basic shape; they are all spherically symmetrical.

#### The radial distribution function:

The square of the radial distribution function describes the probability of finding an a electron a given distance from the nucleus.

In other words, it tells you how far away the fog starts to tapper off, and it tells you whether or not there are pockets of clear sky in the fog.

The radial distribution function describes how far away from the nucleus the orbital extends and the number of nodal surfaces the orbital has.

The radial distribution functions depend on **both n** and l. This means that the number of nodal surfaces and an orbital has and how far that orbital extends from the nucleus depends on the principle quantum number or energy level of the orbital (the 1 in 1s, the 2 in 2s, the 3 in 3s, etc.) and the type of orbital (s vs. p vs. d l = 0, 1, or 2).

Orbitals (n, l, and m <sub>l</sub> )	l is the orbital ty	pe. l=0 15 an 5 Section 2.2.2 , no angular dependence to the
Quantum Numbers	Angular Distribution Function	Radial Distribution Function was function
$n=1$ , $\ell=0$	$\chi(s) = \left(\frac{1}{4\pi}\right)^{\frac{1}{2}}$	$R(1s) = 2\left(\frac{Z}{a_0}\right)^{\frac{3}{2}} e^{(\sigma)} \sim probability$ Falls to 0 of 09
$n = 2$ , $\ell = 0$	$\chi(s) = \left(\frac{1}{4 \pi}\right)^{\frac{1}{2}}$	$R(2s) = \frac{1}{2\sqrt{2}} \left(\frac{Z}{a_0}\right)^{\frac{3}{2}} (2-\sigma) e^{(-\frac{\sigma}{2})}$ probability falls to $R(3s) = \frac{2}{81\sqrt{3}} \left(\frac{Z}{a_0}\right)^{\frac{3}{2}} (27-18\sigma+2\sigma^2) e^{(-\frac{\sigma}{2})}$
$n=$ 3, $\ell=$ 0	$\chi(s) = \left(\frac{1}{4 \pi}\right)^{\frac{1}{2}}$	$R(3s) = \frac{2}{81\sqrt{3}} \left(\frac{Z}{a_0}\right)^{\frac{3}{2}} (27 - 18\sigma + 2\sigma^2) e^{\left(-\frac{\sigma}{3}\right)}$
$n=2,\ell=1$	$\chi(P_x) = \left(\frac{3}{4\pi}\right)^{\frac{1}{2}} \sin \theta \cdot \cos \phi$ $\chi(P_y) = \left(\frac{3}{4\pi}\right)^{\frac{1}{2}} \sin \theta \cdot \sin \phi$	$R(2p) = \frac{1}{2\sqrt{6}} \left(\frac{Z}{a_0}\right)^{\frac{3}{2}} \sigma e^{\left(-\frac{\sigma}{2}\right)}$
g - 7*/o.	$\chi(P_z) = \left(\frac{3}{4\pi}\right)^{\frac{1}{2}} \cos \theta$	

 $\sigma = Zr/a_0$   $a_0 = \varepsilon_0 h^2/(\pi e^2 m_e)$ 

Principles of Mordern Chemistry, 2nd edition. Oxtoby, D.W.; Nachtrieb, N. H. Saunders College Publishing, 1990.

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perode fable 15 shaped by quantum mechanics
The so called families of elements react similarly because, as quantum mechanics tell us, they have similar arrangements of electrons in their highest energy orbitals

Periodic Table of the Elements!
                                                                            7
                                                                               8
                                                                                   9 | 10 | 11 | 12 | 13 | 14 | 15 | 16 | 17
        1=4
                                                   69 70
Tm Yb
                                                                   Ta W
                                                                           Re Os
                                                           103 104 105 106 107 108 109 110 111 112
                                       Cf Es Fm Md No
                                                               Rf Db Sg
                                                                           Bh Hs Mt Ds Rg Cn Nh FI Mc Lv
                                                                                 all halogens are 15° Np5
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